

Background

The PACE (Plankton, Aerosol, Clouds and ocean Ecosystem) mission is a strategic Climate Continuity mission, included in NASA's 2010 plan: "Responding to the Challenge of Climate and Environmental Change: NASA's Plan for a Climate-Centric Architecture for Earth Observations and Applications from Space". On a polar orbit, PACE will make climate-quality global measurements that are essential for understanding ocean biology, biogeochemistry, ecology, aerosol and cloud properties, and determining how the ocean and atmosphere's role in global biogeochemical cycling, ocean ecology and perturbations to Earth's energy balance both affects and is affected by climate change.

With advanced global remote sensing capabilities PACE is expected to: 1. provide high quality observations which will contribute substantially for both basic science and research as well as application and 2. Extend the current time-series of climate quality data to enable detection of long-term trends.

The PACE Science Team 1 (2014-2017)

The PACE science team has been selected in ROSES 2014 and has a three year span. A new science team will be assembled as part of a ROSES 2017 call.

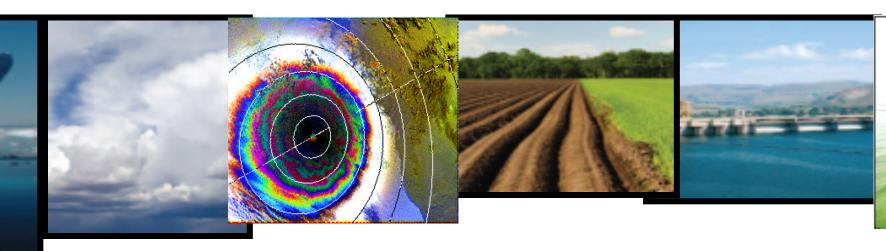
Charge of the first PACE Science Team: collaboratively address inversion for inwater optical properties, and atmospheric characterization, including atmospheric correction.

Science Team goals:

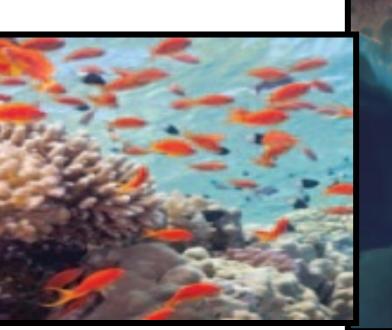
- •Achieve consensus about, and develop community-endorsed paths forward for the PACE sensor(s) for the full spectrum of IOP and atmospheric measurements, algorithms and retrievals.
- Identify gaps in knowledge, research and in technologies that should be filled to maximize the utility of PACE and contribute to it.

In addition, the ST is called upon to provide advise on imager and polarimeter specifications (e.g. we had a series of webminars to learn about new/available technologies, their attributes and the science they will facilitate).

While th members of the team each have been funded to accomplish *specific* science goals, here we highlight some of the team's collaborative accomplishments from its first year.





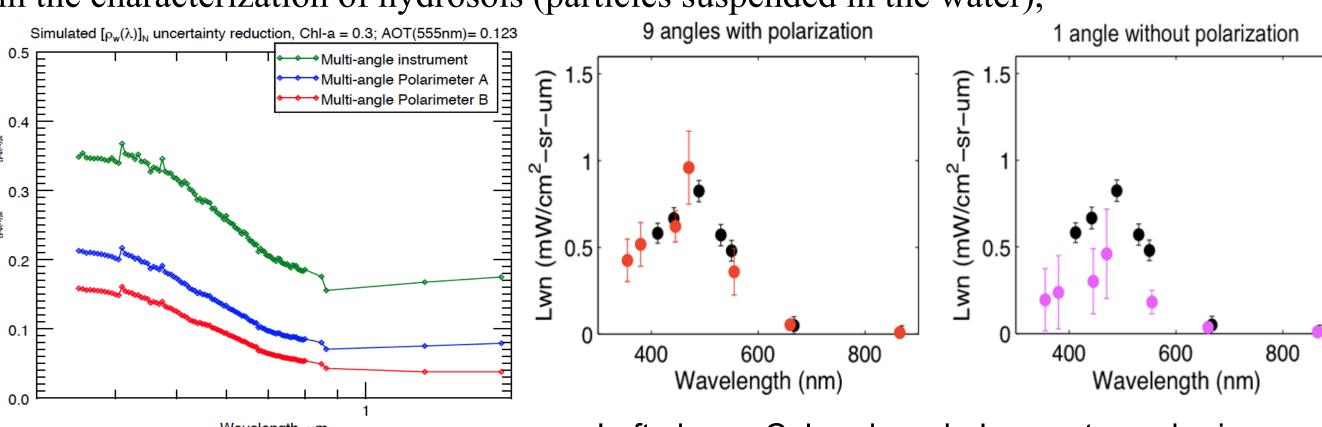


First year accomplishments

Polarimetry Report (Lead: L. Remer)

The PACE Science Team strongly endorses the inclusion of a polarimeter in the PACE mission. This consensus opinion is based on the following three reasons:

- (1) There are specific atmospheric products related to both clouds and aerosols that can only be obtained with a multi-angle polarimeter
- (2) Multi-angle polarimetry will reduce uncertainty in derivation of normalized water leaving radiance and provide more accurate atmospheric correction. (See figures.)
- (3) Preliminary theoretical and observational studies that show that polarimetry can aid in the characterization of hydrosols (particles suspended in the water),



Ratio of uncertainty in normalized water leaving reflectance ($[\rho_w(\lambda)]_N$) between theoretical instruments having multi-angle and polarimetric capability and those measuring only radiometry at one view angle per target.

Left above. Colored symbols are atmospheric correction results derived from high altitude aircraft observations at 9 angles; 7 wavelengths; and with polarization. Right above. Colored symbols are results derived without angular and polarization information. In both, black symbols are measurements from SeaPRISM near sea surface.

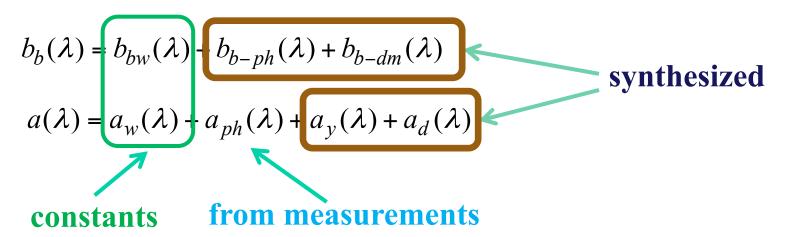
Synthetic hyper-spectral IOP-AOP dataset (Lead: Z. Lee)

•Objective: Create a free-of-measurement-error hyperspectral (350-800 nm, 5 nm resolution) dataset for algorithm test and evaluation.

•General rules:

- Representative to commonly encountered waters.
- Consistent with up-to-date knowledge.

•Approach: create an IOP dataset, and feed it to Hydrolight to generate corresponding AOP spectra. $b_{k}(\lambda) = b_{k}(\lambda) + b_{k}(\lambda) + b_{k}(\lambda)$



•720 a_{ph} spectra (from >4000 SeaBASS spectra): 12 "groups" (separated by a_{ph} (440) value), and ~3 a_{ph} shapes within each "group".

$$a_{y}(\lambda) = a_{y}(440)e^{-S_{y}(\lambda - 440)}; \ a_{y}(440) = p_{1} \times a_{ph}(440)$$

$$a_{d}(\lambda) = a_{d}(440)e^{-S_{d}(\lambda - 440)}; \ a_{d}(440) = p_{2} \times a_{ph}(440)$$

$$b_{b-ph}(\lambda) = 0.01 \times \left(c_{ph}(\lambda) - a_{ph}(\lambda)\right)$$

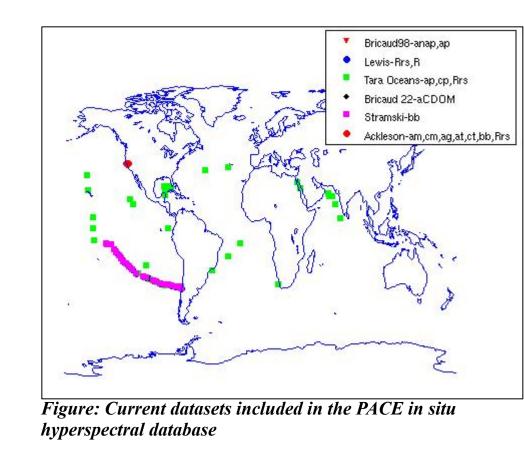
$$b_{b-dm}(\lambda) = p_{5} \times a_{ph}(440) \times \left(\frac{440}{\lambda}\right)^{p_{6}}$$

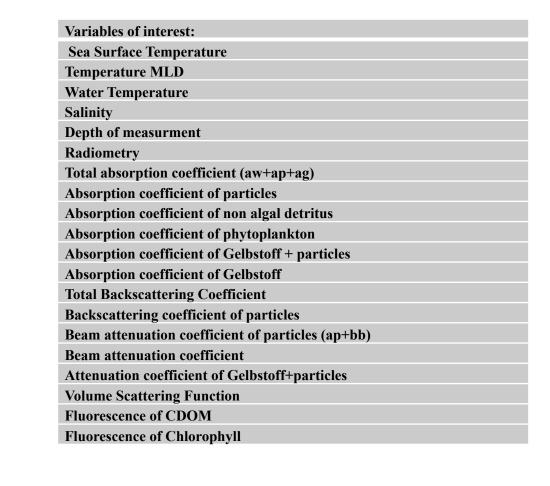
$$c_{ph}(\lambda) = p_{3} \times a_{ph}(440) \times \left(\frac{440}{\lambda}\right)^{p_{4}}$$

•Value of p_1 , S_y , p_2 , S_d , p_3 , p_4 , p_5 , and p_6 are determined semi-randomly (within ranges from field observations) as in IOCCG Report #5.

In-situ dataset (Lead C. Rousseaux)

• **Objective:** To create a high quality, diverse and extensive database of existing hyperspectral inherent and apparent optical properties (IOP/AOP) data both polarized and non-polarized) for the validation of remote sensing products and the development of algorithms for the PACE mission





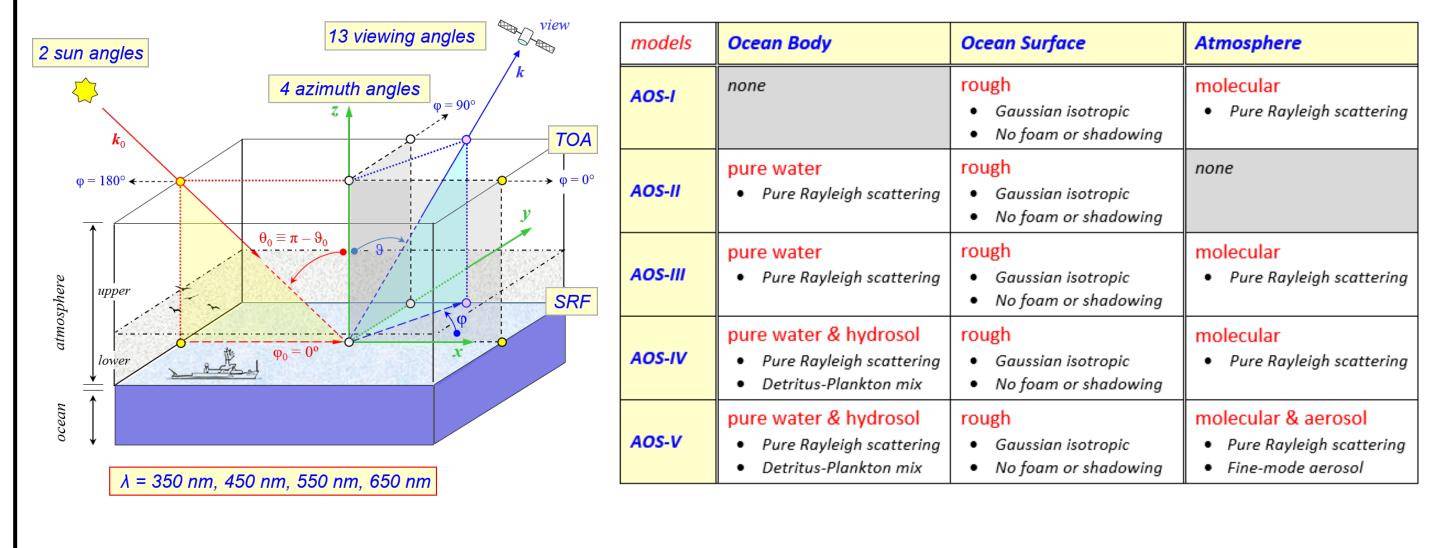
•Outcome: A database of hyperspectral IOP and AOP data covering a range of ecosystems to be used for the validation of PACE products and algorithms. This database will be published with every contributor as co-author in ~2017. Please contact Cecile S. Rousseaux (Cecile.S.Rousseaux@nasa.gov) and Emmanuel Boss (Emmanuel.Boss@maine.edu) for to contribute your data to this database or for any questions, recommendations or comments.

Pre formulation,

conceptual studies

Intercomparision of polarized RT codes (Lead: J. Chowdhary)

- **Objective:** The high measurement accuracy (0.2%-0.5%) of modern polarimeters deployed over oceans requires matching improvements in validating the simulation accuracy for polarized light scattered in atmosphere-ocean system (AOS) models
- Set-up: 5 AOS models, 104 scattering geometries, 2 altitudes, 4 wavelengths:



• **Results:** Using 3 independently written RT codes that were based on 2 different methods, we obtained thus far for AOS-I to AOS-III agreement within 10⁻⁶ for Stokes parameters I, Q, U. This is 1-2 orders of magnitude better than previous comparisons, and leads to an <0.1% accuracy in degree of linear polarization.

At NASA centers:

preliminary Design and development analysis

Launch

 2010
 2011
 2012
 2013
 2014
 2015
 2016
 2017
 2018
 2019
 2020
 2021
 2022
 2023

formulation,

Science teams:

PACE-SDT

PACE-ST-1, atmospheric-correction and inversions for IOPs

PACE-ST-2, Cal/Val planning

PACE-ST-3, pre-launch

NASA

Open science calls:

Vicarious calibration system design

Vicarious calibration system build